Thomas Montandon Elsa M. Teixeira Adele Poudou Vivian Poulin

arXiv:2505.20193



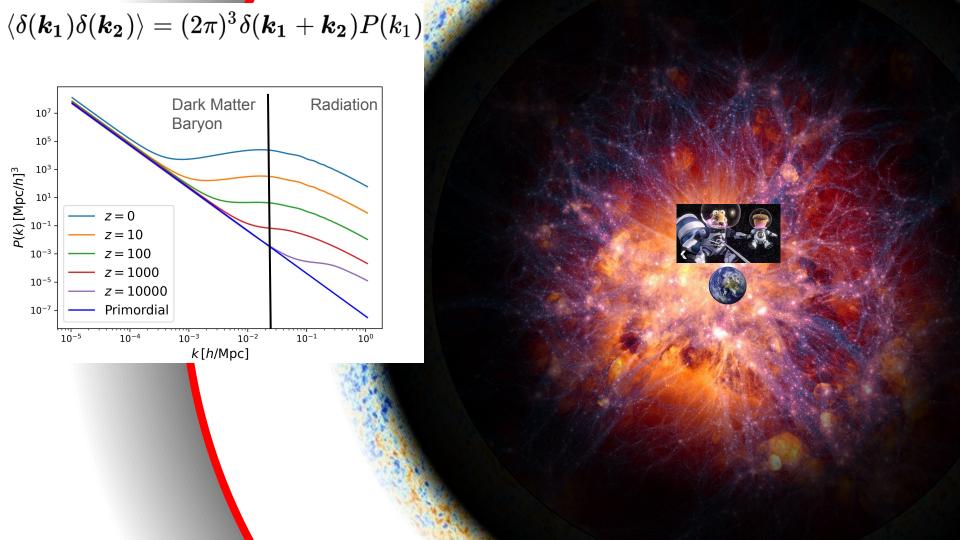










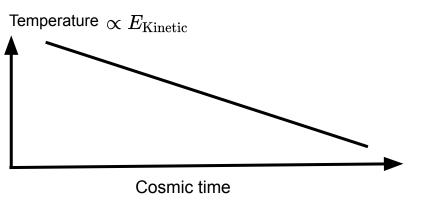


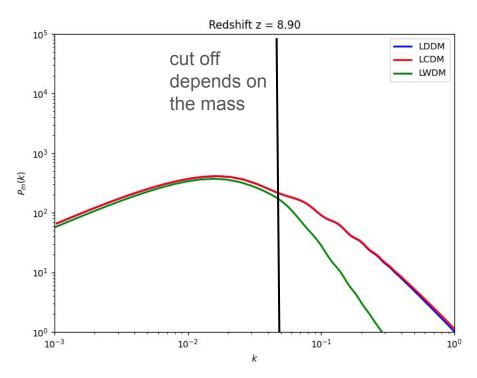
What?

 $\circ$  cold:  $E_{
m kinetic} \ll m$ 

 $\circ$  hot:  $E_{
m kinetic}\gg m$ 

o warm:





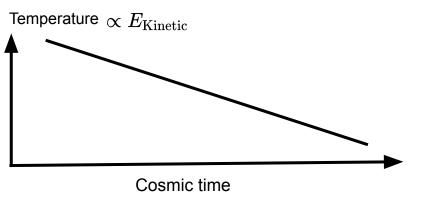
 $E_{
m kinetic}\gg m$   $\longrightarrow$   $E_{
m kinetic}\sim m$   $\longrightarrow$   $E_{
m kinetic}\ll m$ 

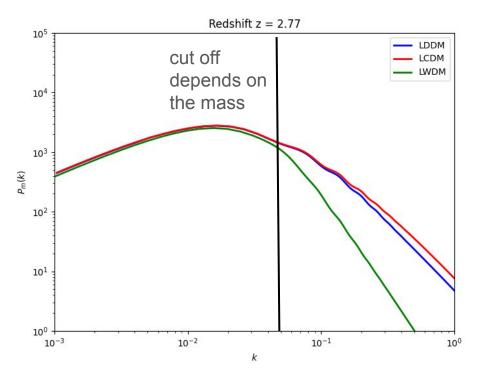
What?

 $\circ$  cold:  $E_{
m kinetic} \ll m$ 

 $\circ$  hot:  $E_{
m kinetic}\gg m$ 

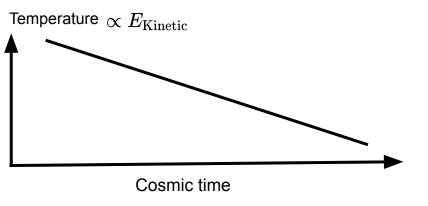
o warm:

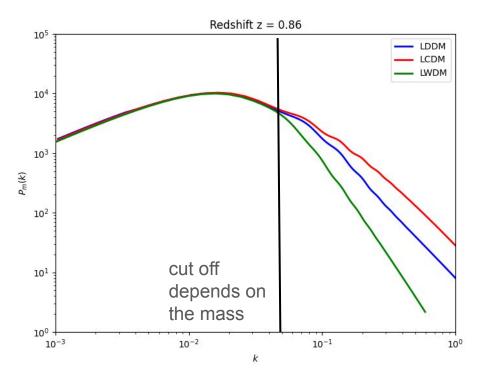




 $E_{
m kinetic}\gg m$   $\longrightarrow$   $E_{
m kinetic}\sim m$   $\longrightarrow$   $E_{
m kinetic}\ll m$ 

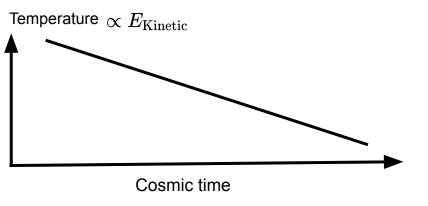
- What?
  - $\circ$  cold:  $E_{
    m kinetic} \ll m$
  - $\circ$  hot:  $E_{
    m kinetic}\gg m$
  - o warm:

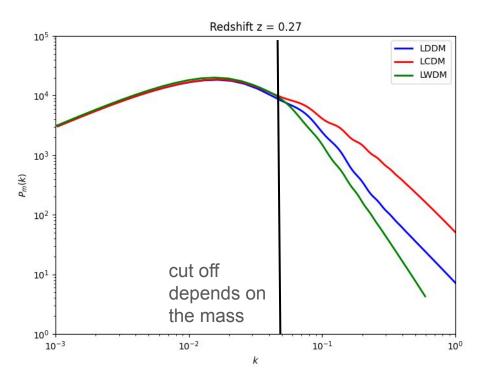




$$E_{
m kinetic}\gg m$$
  $\longrightarrow$   $E_{
m kinetic}\sim m$   $\longrightarrow$   $E_{
m kinetic}\ll m$ 

- What?
  - $\circ$  cold:  $E_{
    m kinetic} \ll m$
  - $\circ$  hot:  $E_{
    m kinetic}\gg m$
  - o warm:





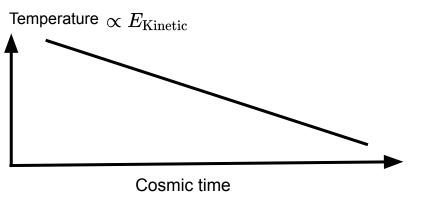
$$E_{
m kinetic}\gg m$$
  $\longrightarrow$   $E_{
m kinetic}\sim m$   $\longrightarrow$   $E_{
m kinetic}\ll m$ 

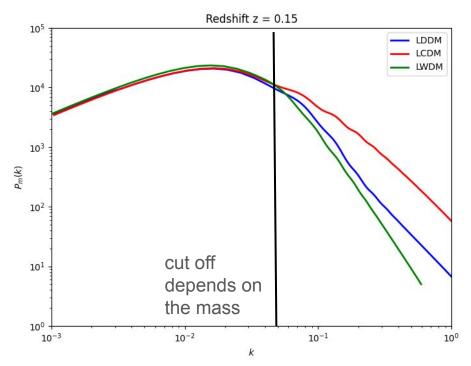
What?

 $\circ$  cold:  $E_{
m kinetic} \ll m$ 

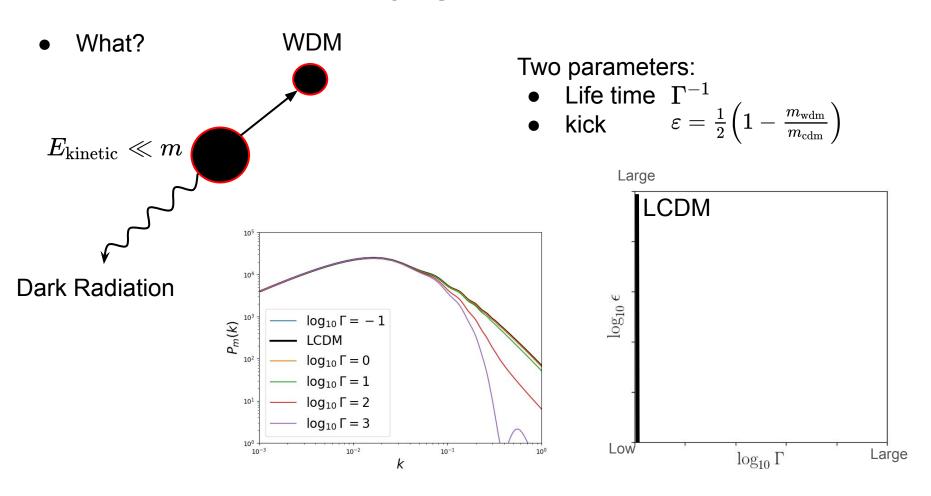
 $\circ$  hot:  $E_{
m kinetic}\gg m$ 

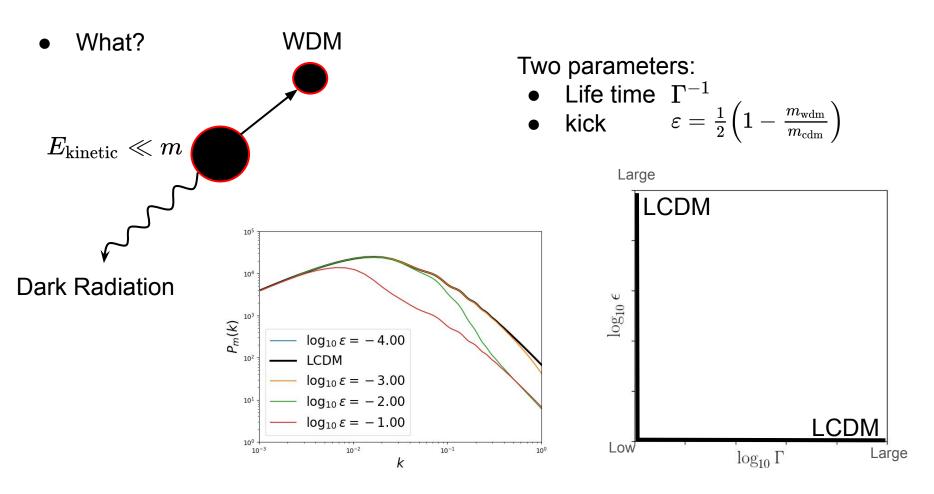
o warm:

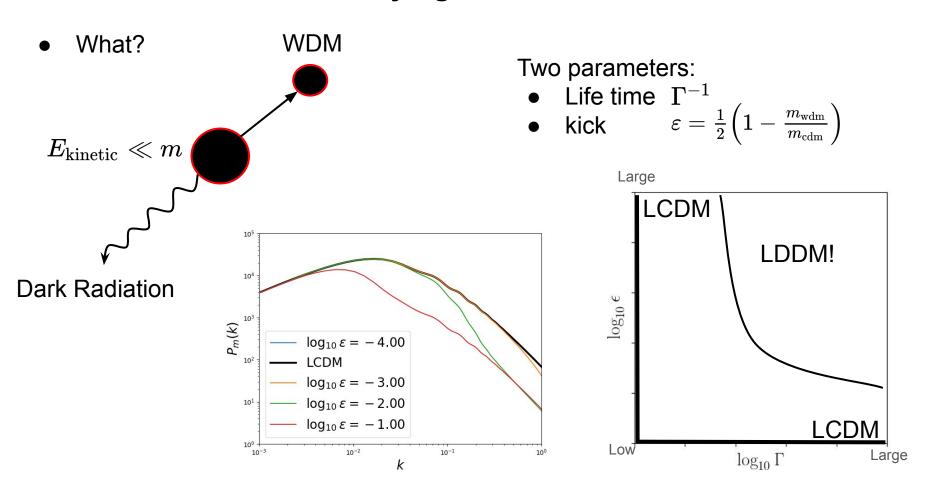




 $E_{
m kinetic}\gg m$   $\longrightarrow$   $E_{
m kinetic}\sim m$   $\longrightarrow$   $E_{
m kinetic}\ll m$ 





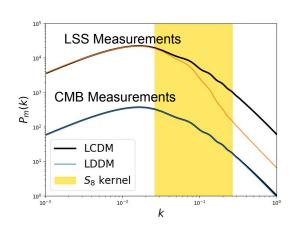


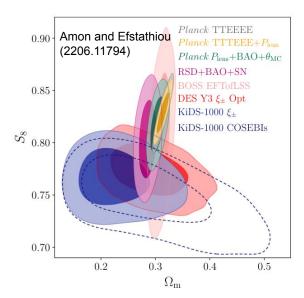
# Sigma-8 tension

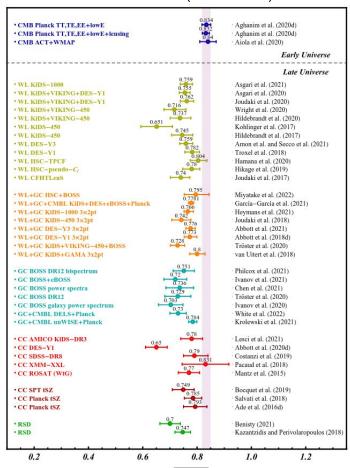
#### Abdalla et al (2203.06142)

Why?
$$\sigma_R^2 = \int dk rac{k^2 P_m(k)}{2\pi^2} W^2(kR)$$

 $R=8h^{-1}{
m Mpc}\sim{
m Galaxy}$  clustering scale







# Sigma-8 tension

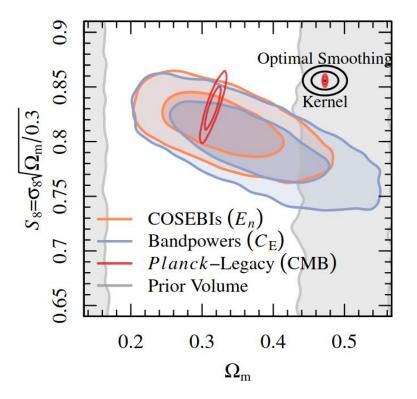
Why?

$$\sigma_R^2=\int dk rac{k^2 P_m(k)}{2\pi^2} W^2(kR)$$

 $R=8h^{-1}{
m Mpc}\sim{
m Galaxy}$  clustering scale



The KiDS Collaboration: 2503.19441



# Sigma-8 tension

Why?

$$\sigma_R^2 = \int dk rac{k^2 P_m(k)}{2\pi^2} W^2(kR)$$

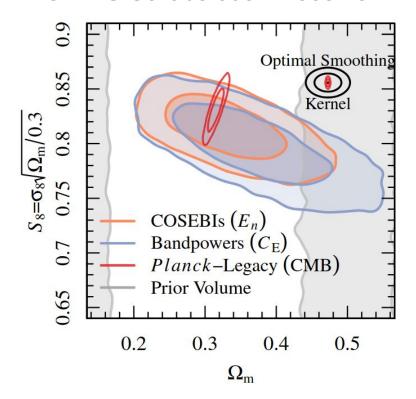
 $R=8h^{-1}{
m Mpc}\sim{
m Galaxy}$  clustering scale



We can still test nature of DM!

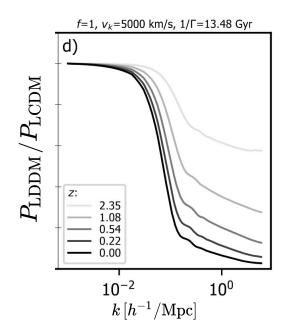


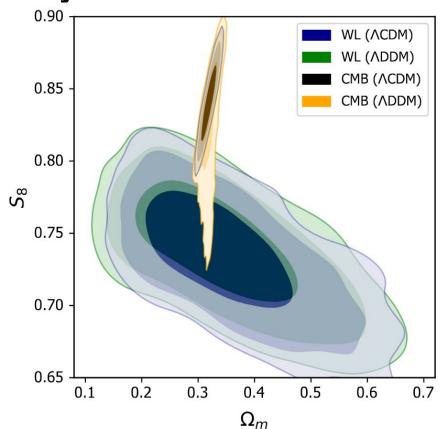
The KiDS Collaboration: 2503.19441



# Decaying Dark Matter Bayesian analysis

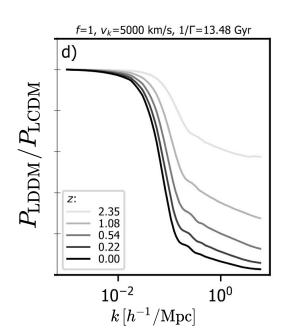
- How?
- For LSS we need Non Linearity!
   J.Bucko, A.Schneider et al. 2307.03222
  - DDM simulation with PKDGRAV3
  - Neural Network fit of P(k)

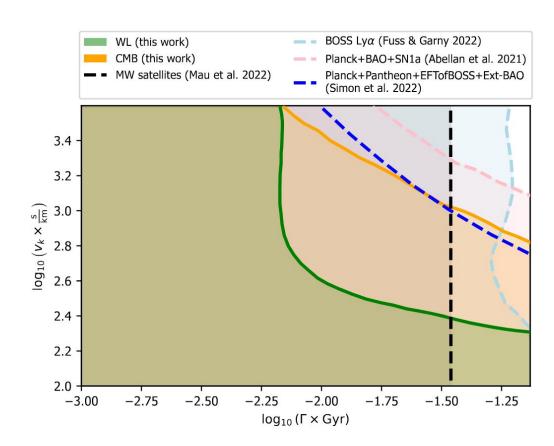




# Decaying Dark Matter Bayesian analysis

- How?
- For LSS we need Non Linearity!
   J.Bucko, A.Schneider et al. 2307.03222
  - DDM simulation with PKDGRAV3
  - Neural Network fit of P(k)



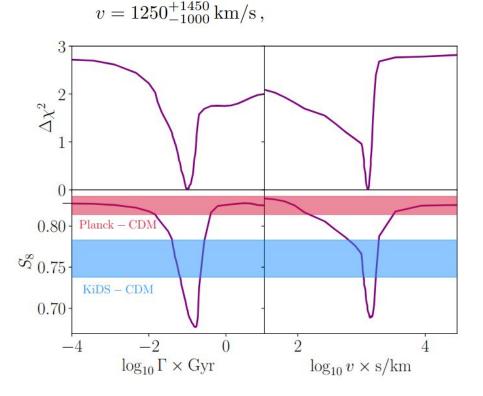


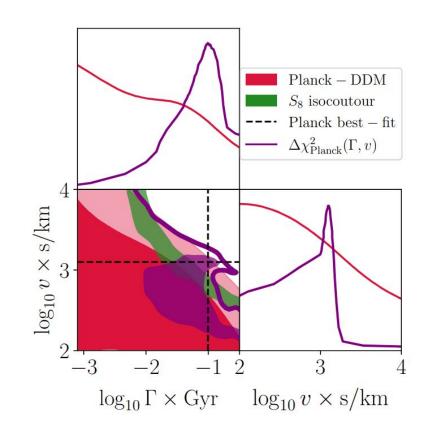
$$\Gamma = 0.10^{+0.17}_{-0.05} \,\text{Gyr}^{-1},$$

$$t_{1/2}^{\text{DDM}} = 6.93^{+7.88}_{-2.85} \,\text{Gyr},$$

$$N_0^{\text{DDM}}/N_{\text{ini}} = 0.25_{-0.16}^{+0.27},$$

Planck: There is something here!



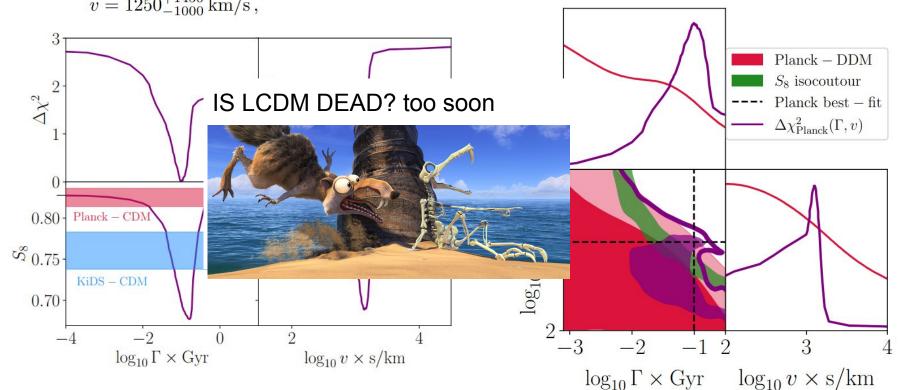


 $\Gamma = 0.10^{+0.17}_{-0.05} \,\mathrm{Gyr}^{-1}$  $t_{1/2}^{\text{DDM}} = 6.93_{-2.85}^{+7.88} \,\text{Gyr}\,,$ 

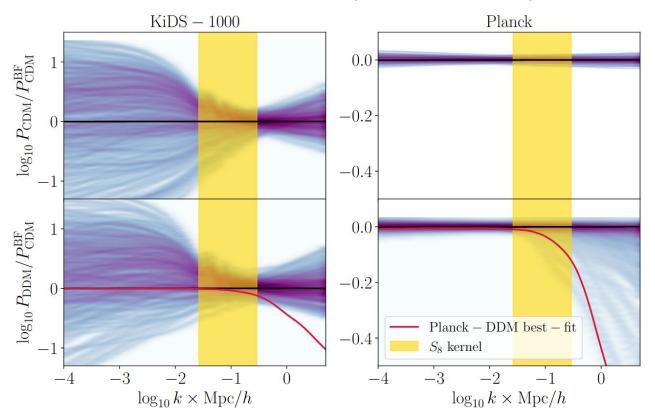
 $N_0^{\text{DDM}}/N_{\text{ini}} = 0.25_{-0.16}^{+0.27}$ ,

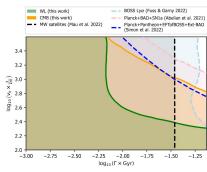
 $v = 1250^{+1450}_{-1000} \,\mathrm{km/s}$ 

Planck: There is something here!

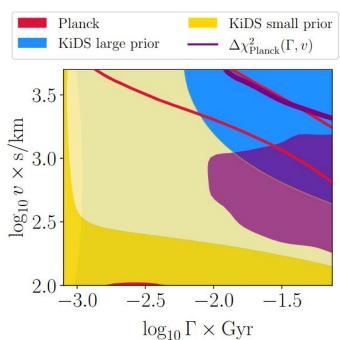


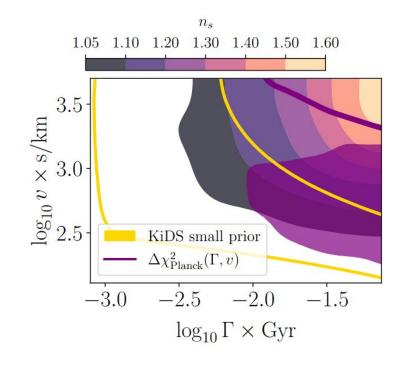
- Planck: There is something here!
- What KiDS really measures: not just S8!



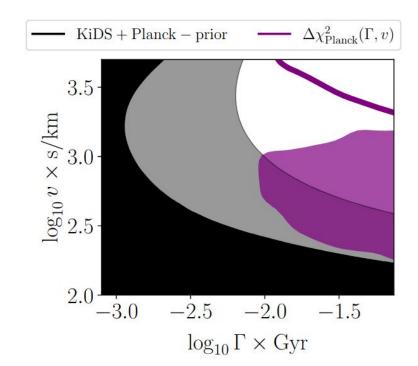


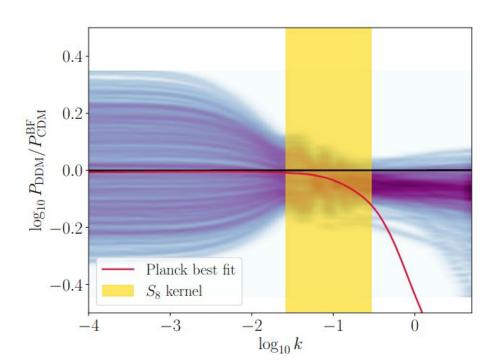
- Planck: There is something here!
- What KiDS really measures: not just S8!
- KiDS: a prior issue...





- Planck: There is something here!
- What KiDS really measures: not just S8!
- KiDS: a prior issue...
- KiDS and Planck

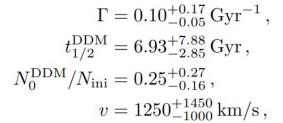


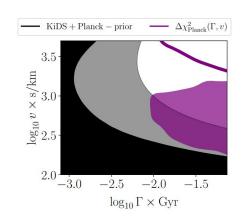


## Conclusion

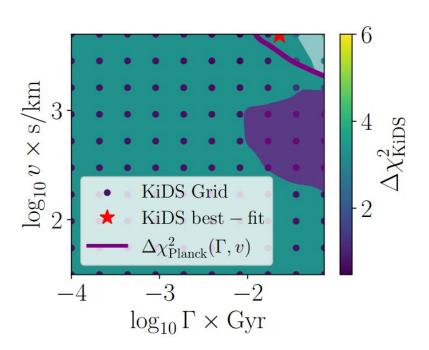
- Planck-2018 + BOSS-BAO + Pantheon-Plus
  - Subject to volume effect
  - Best-fit S8 compatible with KiDS
- KiDS-1000
  - Cannot be reduced to its S8 measure (eg S8-prior)
  - Subject to prior effect
- KiDS + Planck-informed prior





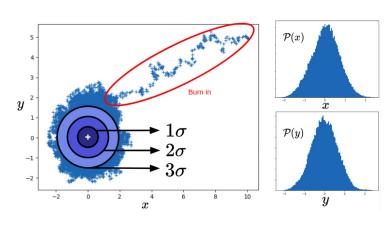


KiDS profile likelihood



# **Bayesian vs Frequentist**

#### Bayesian: density of points



Profile likelihood: prior independent!

### A.Nygaard et al 2308.06379

